

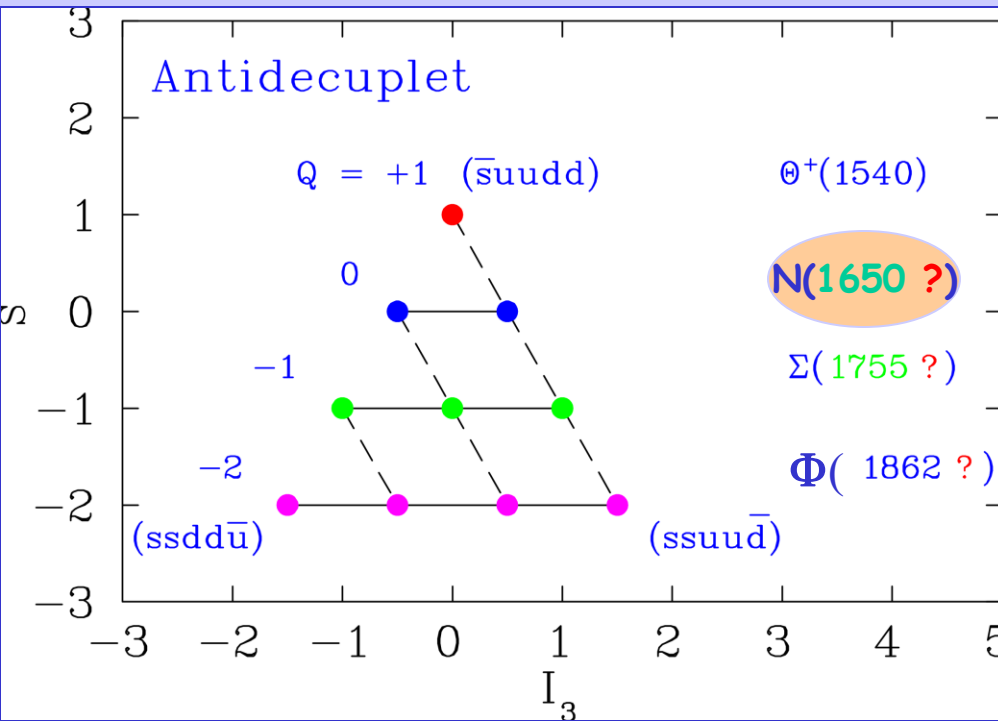
N^* as a Flavor Partner of the Θ^+ . Where are we now?

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Based on work in collaboration with
R. Arndt, Ya. Azimov, W. Briscoe, M. Polyakov, R. Workman

- Antidecuplet
- Is $N^* = N(1710)$?
- How to search for alternatives ?
Modified PWA
- Theoretical expectations
- Experimental evidence for N^*
- Summary

Tentative unitary Antidecuplet with Θ^+



- **GMO**: equidistant, expected $\delta m(\sigma) = (M_\Phi - M_\Theta)/3$ depends on σ -term
 =107 MeV at $\sigma = 67$ MeV [SAID]
 =180 MeV at $\sigma = 45$ MeV [Karlsruhe]
- δm agrees with the *GW SAID* σ -term, if $M_\Theta = 1540$ MeV and $M_\Phi = 1862$ MeV
- Mixing tends to shift **GMO** masses for N^* and Σ^* stronger, than for Θ and Φ

N(1710) - What was known

[S. Eidelman *et al* (PDG) Phys Lett B **592**, 1 (2004)]

χ SA	DPP97	1710 (inp)	~40 (est)
PWA-BW	Ref	Mass(MeV)	Width(MeV)
	KH79	1723± 9	120± 15
	CMU80	1700±50	90± 30
	KSU92	1717±28	480±230

GW04

not seen ←

No BW, No pole, No Sp

PWA-Pole

Re(MeV) -2xIm(MeV)

CMU80	1690±20	80± 20
CMU90	1698	88
KH93	1690	200

(Sp)

GW04

not seen

- It would be more natural for the same unitary multiplet (with Θ^+ and N^*) to have comparable widths
- The spread of Γ , separated by PDG, is very large

Standard Resonances in Standard PWA

[R. Arndt, W. Briscoe, IS, R. Workman, M. Pavan, Phys Rev C 69, 035213 (2004)]

- One of the most convincing ways to study N^* s and Δ^* s is πN PWA
- Standard PWA reveals only **wide** resonances, but not too wide ($\Gamma < 500$ MeV)
- PWA (by construction) tends to miss **narrow** resonances with $\Gamma < 30$ MeV

Narrow Resonances in PWA

[R. Arndt, Ya. Azimov, M. Polyakov, IS, R. Workman, Phys Rev C **69**, 035208 (2004)]

- We **assume** the existence of a Res and **refit** over the whole database
- Insertion of narrow resonances in PWA for
elastic case: $e^{2i\delta} \Rightarrow e^{2i\delta}_R e^{2i\delta}_B$
$$e^{2i\delta}_R = (M_R - W + i\Gamma_R/2)/(M_R - W - i\Gamma_R/2)$$

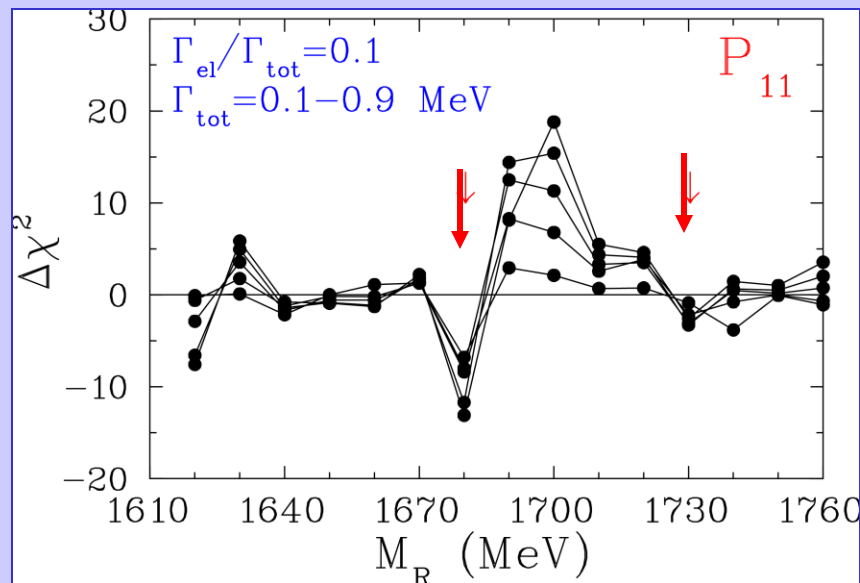
inelastic case: $\eta e^{2i\delta} \Rightarrow \langle a|S|a \rangle = r_a A(W) e^{2i\delta}_R + (1 - r_a) B(W)$
$$r_a = BR(R \rightarrow a) \quad |A(M_R)| = 1 \quad \Sigma r_a = 1$$

$$\eta \leq 1 \Rightarrow r_a |A(W)| + (1 - r_a) |B(W)| \leq 1$$
- How does this insertion change χ^2 ?
(Will it decrease?)

Modified πN PWA

[R. Arndt, Ya. Azimov, M. Polyakov, IS, R. Workman, Phys Rev C **69**, 035208 (2004)]

- $\Delta\chi^2$ due to insertion of a resonance into P_{11} ($J^P = \frac{1}{2}^+$)



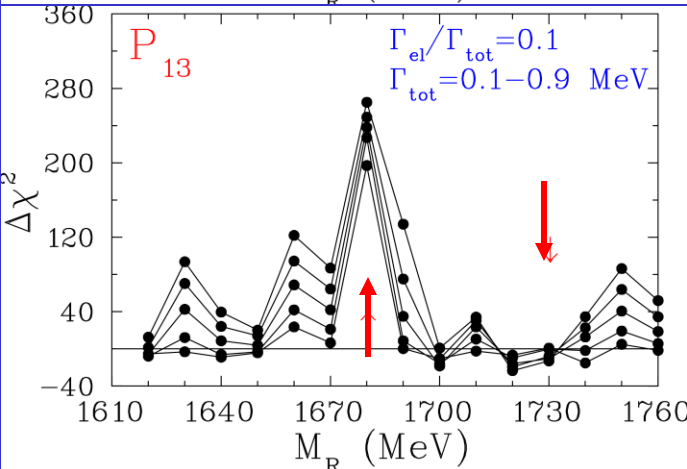
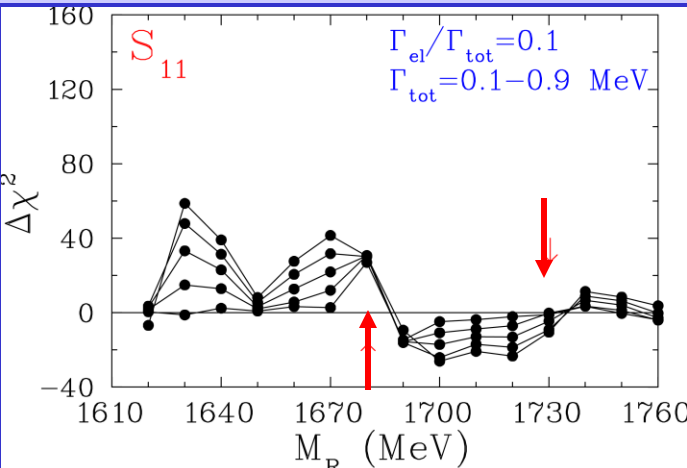
- At $|M_R - W| \gg \Gamma_R$, Res contributes $\sim \Gamma_{el}/(M_R - W)$
- Two candidates:

$M_R = 1680 \text{ MeV}$	1730 MeV
$\Gamma_{\pi N} < 0.5 \text{ MeV}$	$< 0.3 \text{ MeV}$
- The procedure is less sensitive to Γ_{tot}

Modified PWA

- Refitting
 - Worse description
 - ⇒ a **Res** with corresponding M and Γ is not supported
 - Better description
 - ⇒ a **Res** may exist
 - ⇒ effect can be due to various corrections (*eg*, thresholds)
 - ⇒ both possibilities can contribute
- Some additional checks are necessary
- A true **Res** should provide the effect only in a particular PW
 - While **NonRes** source may show similar effects in various PWs

Check other Partial Waves



- $\Delta\chi^2$ due to insertion of a Res into S_{11} ($J^P = 1/2^-$)

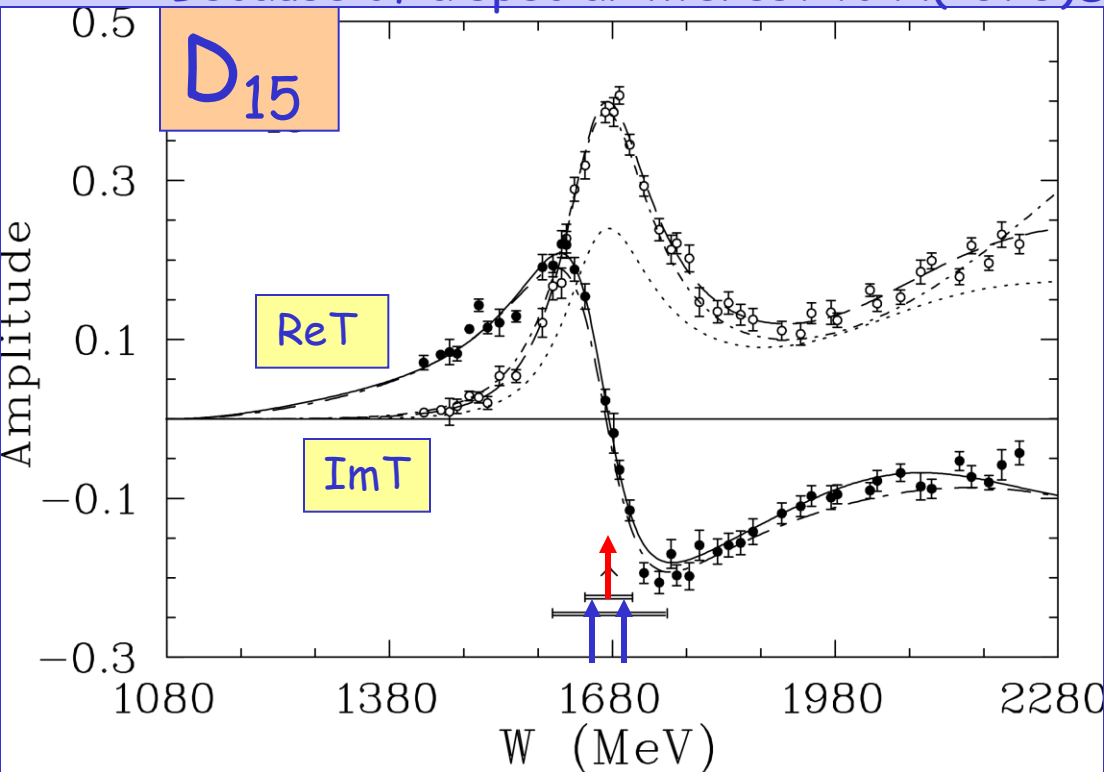
- $\Delta\chi^2$ due to insertion of a Res into P_{13} ($J^P = 3/2^+$)

- No effects at $M = 1680$ MeV and possible (small) effects at $M = 1730$ MeV

D_{15} within πN PWA

[R. Arndt, W. Briscoe, IS, R. Workman, M. Pavan, Phys Rev C **69**, 035213 (2004)]

- Because of a special interest to $N(1675)D_{15}$, let us check the situation



$\text{Im}T - T^*T < \text{Im}T$ [unitarity limit]

- $N(1675)D_{15}$ [SAID]
 - BW: $M = 1676.2 \pm 0.6$ MeV
 - $\Gamma/2 = 75.9 \pm 1.5$ MeV
 - $X = 0.400 \pm 0.002$
 - Pole: $1659 - i73$ MeV

- $N(1675)D_{15}$ has a standard BW and pole
- It is unnatural to have two nearby resonances with the same quantum numbers

N(1675)D₁₅ - What is known

[S. Eidelman *et al* (PDG) Phys Lett B **592**, 1 (2004)]

PWA-BW	Ref	Mass(MeV)	Width(MeV)	$\Gamma_{\pi N}/\Gamma$
	KH79	1679± 8	120±15	0.38±0.03
	CMU80	1675±10	160±20	0.38±0.05
	KSU92	1676± 2	159± 7	0.47±0.02
	GW04	1676.2±0.6	151.8±3.0	0.400±0.002
PWA-Pole		Re(MeV)	-2xIm(MeV)	
	CMU80	1660±10	140±10	
	KH93	1656	126	
	GW04	1659	146	

N(1675)D₁₅ - What is known Other Channels

- $\Gamma_{\pi N}/\Gamma = 0.40 \pm 0.05$ [PDG]
 - $\Gamma_{\pi\Delta}/\Gamma = 0.496 \pm 0.003$ [KSU92]
 - $\Gamma_{\rho N}/\Gamma = 0.03 \pm 0.02$ [KSU92]
 - $\Gamma_{K\Lambda}/\Gamma = 0.036$ [Ruth80]
 - $\Gamma_{\eta N}/\Gamma = 0.00 \pm 0.01$ [CMU00]
-
- There is really no room for **large BR** of N(1675)D₁₅ into other decay channels

Conclusion from Modified πN PWA for S- and P-waves

[R. Arndt, Ya. Azimov, M. Polyakov, IS, R. Workman, Phys Rev C **69**, 035208(2004)]

- **1680 MeV** - only one partial wave (P_{11}) reveals the effect: support to the resonance, $\Gamma_{\pi N} < 0.5$ MeV
- **1730 MeV** - P_{11} may also reveal a resonance with $\Gamma_{\pi N} < 0.3$ MeV but differently: resonance is still possible, if accompanied by different corrections
- The Res at **1730 MeV** may appear in P_{13} or S_{11} (less probable), if accompanied by different corrections [eg, thresholds: $N\omega(1720)$, $N\rho(1710)$?, $K\Sigma(1685)$]
- The rest of partial waves (D_{15} , etc) do not support narrow states

Θ^+ Flavor Partner, $N^*(J^P = \frac{1}{2}^+)$

[R. Arndt, Ya. Azimov, M. Polyakov, IS, R. Workman, Phys Rev C **69**, 035208 (2004)]

- If $\Gamma_{\Theta} \leq 1 \text{ MeV}$, then expected structure for decays of the Θ -partner N^* looks as follows:
 - $\Gamma(N^* \rightarrow \pi\Delta) \sim 6 \text{ MeV}$ [forbidden for $\overline{10}$, open due to $\overline{10}$ -8 mixing]
 - $\Gamma(N^* \rightarrow \eta N) \sim 0.5 - 2 \text{ MeV}$
 - $\Gamma(N^* \rightarrow K\Lambda) \sim 0.5 - 1.5 \text{ MeV}$
 - $\Gamma(N^* \rightarrow \pi N) \sim 0.3 - 0.5 \text{ MeV}$ [non-trivial cancellation due to mixing is required]
 - $\Gamma(N^* \rightarrow \pi\pi N)$ [out of $\pi\Delta$] ?
 - $\Gamma(N^* \rightarrow K\Sigma)$ is small ?
 - $\Gamma(N^* \rightarrow \text{all}) \sim 10 \text{ MeV}$ [$\Gamma_{\pi N}/\Gamma_{\text{tot}} < 10 \%$]

Ratio of modes πN and ηN is sensitive to the mixing

Preliminary Evidences for Narrow State(s) of $M \sim 1700$ MeV

- **GRAAL**: $\gamma n \rightarrow \eta n$, $K^0 \Lambda$, and $K^+ \Sigma^-$
- **CB-ELSA**: $\gamma n \rightarrow \eta n$
- **JLab Hall A**: $H(e, e' \pi^+) X^0$

- **STAR**: $AuAu \rightarrow \Lambda K_s$
- **COSY-TOF**: $pp \rightarrow \Lambda K^+ p$

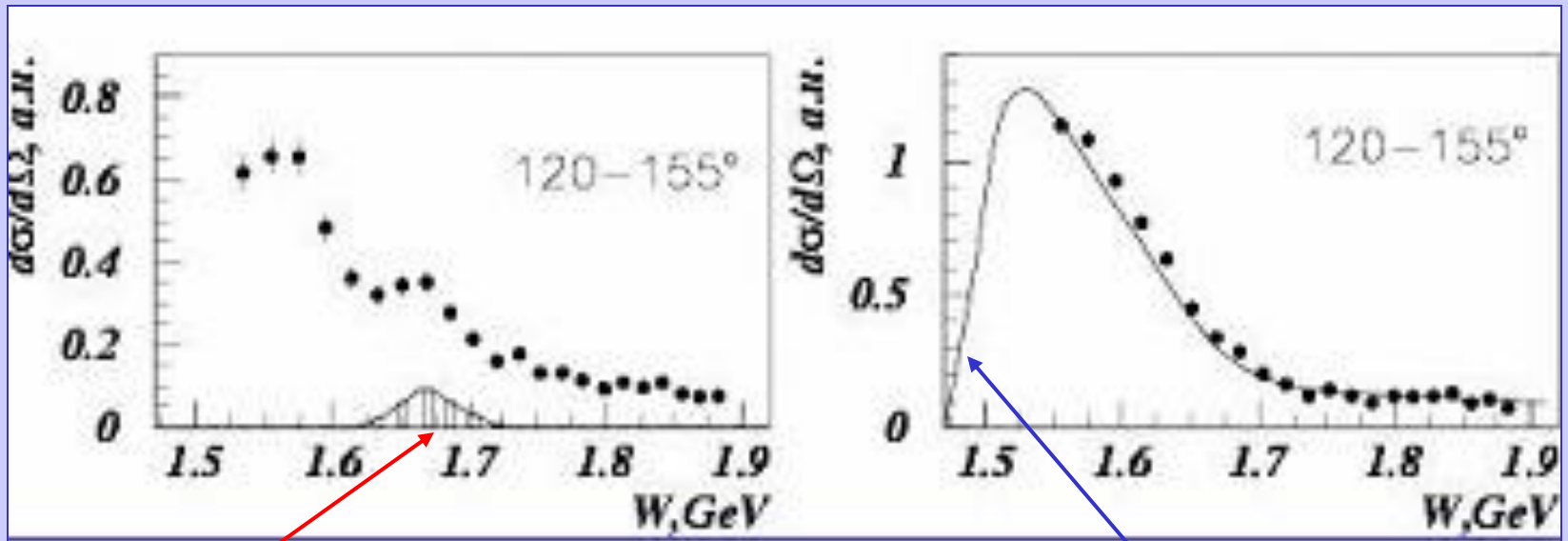
- **ITEP**: $\pi^- p \rightarrow \pi^- p$ and $K^0 \Lambda$ [in preparation]

GRAAL [V. Kuznetsov, hep-ex/0409032, NSTAR 2004, March 2004]

$\gamma n \rightarrow \eta n$ vs $\gamma p \rightarrow \eta p$

$\gamma n \rightarrow \eta n$

$\gamma p \rightarrow \eta p$



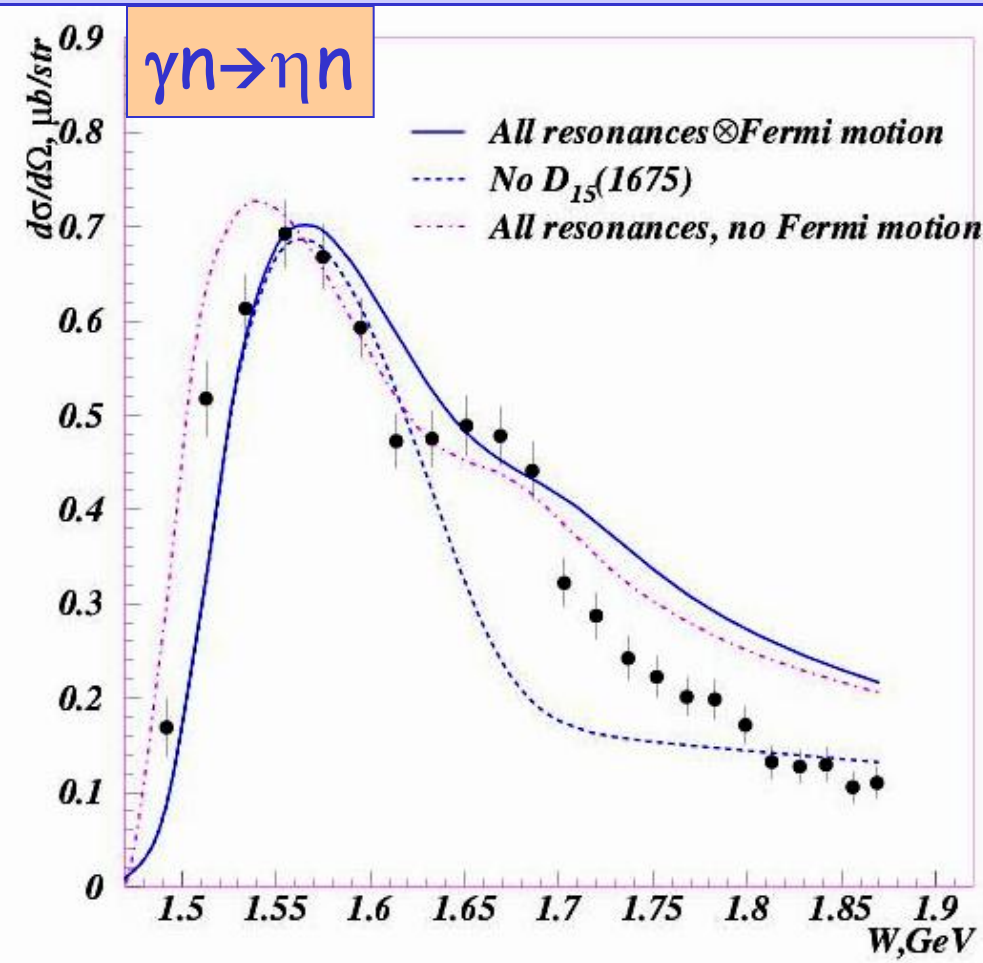
N(1670)

SAID PWA

- For $\bar{10}$, $\sigma(n) \gg \sigma(p)$ [M. Polyakov, A. Ratke, Eur Phys J A **18**, 691 (2003)]
- Fermi motion for n-target is a problem

MAID about GRAAL Observation

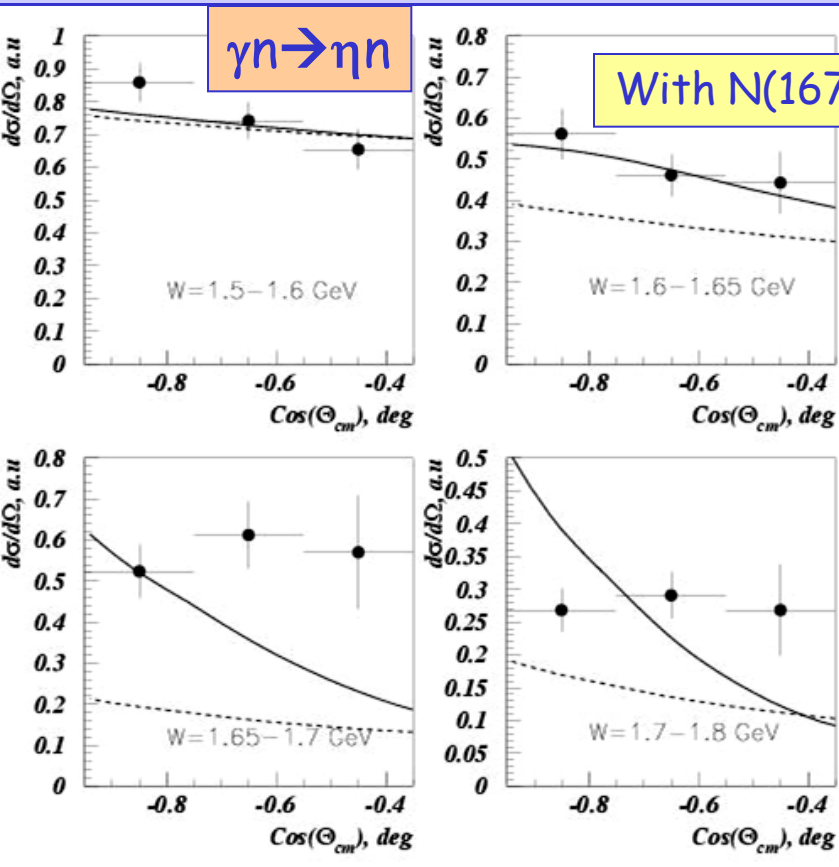
[V. Kuznetsov, hep-ex/0601002, NSTAR 2005, Oct 2005]



- MAID2000 demonstrates a shoulder structure near $N(1675)D_{15}$
- MAID2000 claims to reproduce the rise in the ratio of the neutron/proton cross sections
- However, the experimental structure looks more narrow

MAID about GRAAL Observation

[V. Kuznetsov, hep-ex/0601002, NSTAR 2005, Oct 2005]



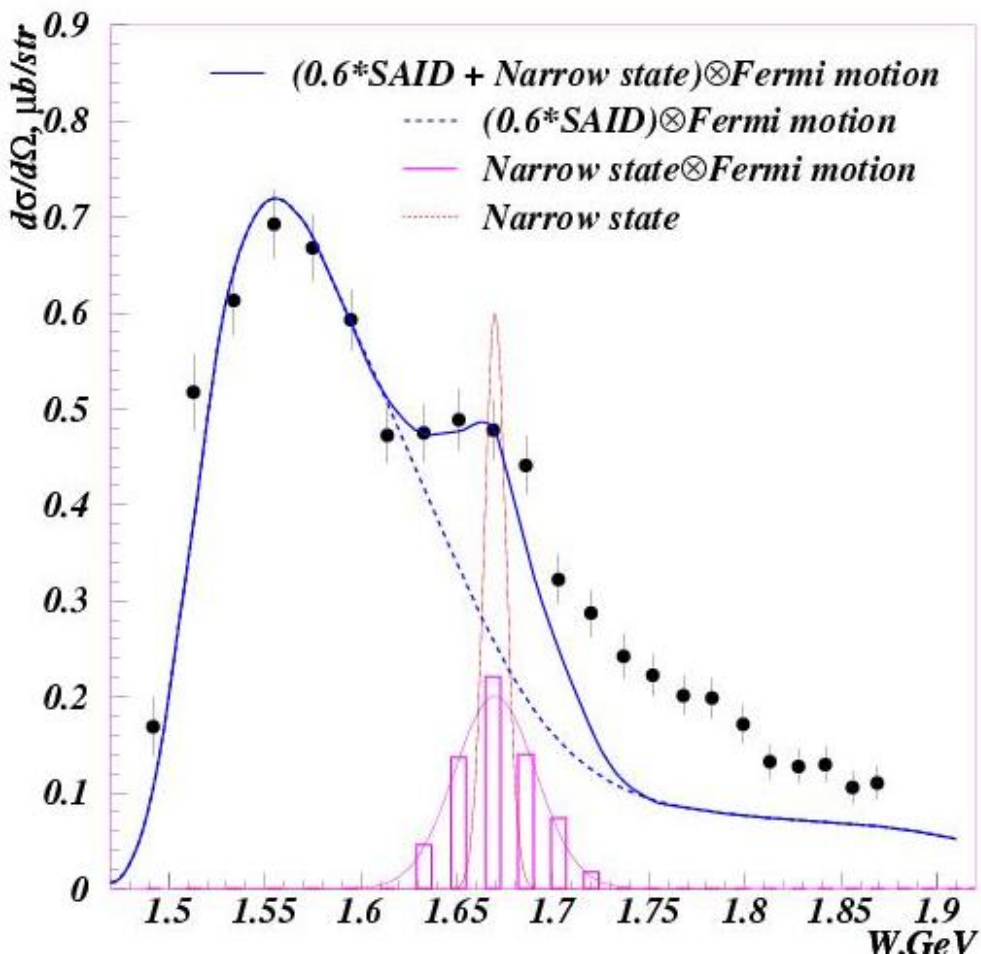
- The N(1675)D₁₅ signature is a strong (incorrect ?) angular dependence in Xsection

No N(1675)D₁₅

- Larger spin steepens Xsection dramatically
- Same problem for Σ behavior

GRAAL [V. Kuznetsov, hep-ex/0601002, NSTAR 2005, Oct 2005]

Very preliminary: $\gamma n \rightarrow \eta n$



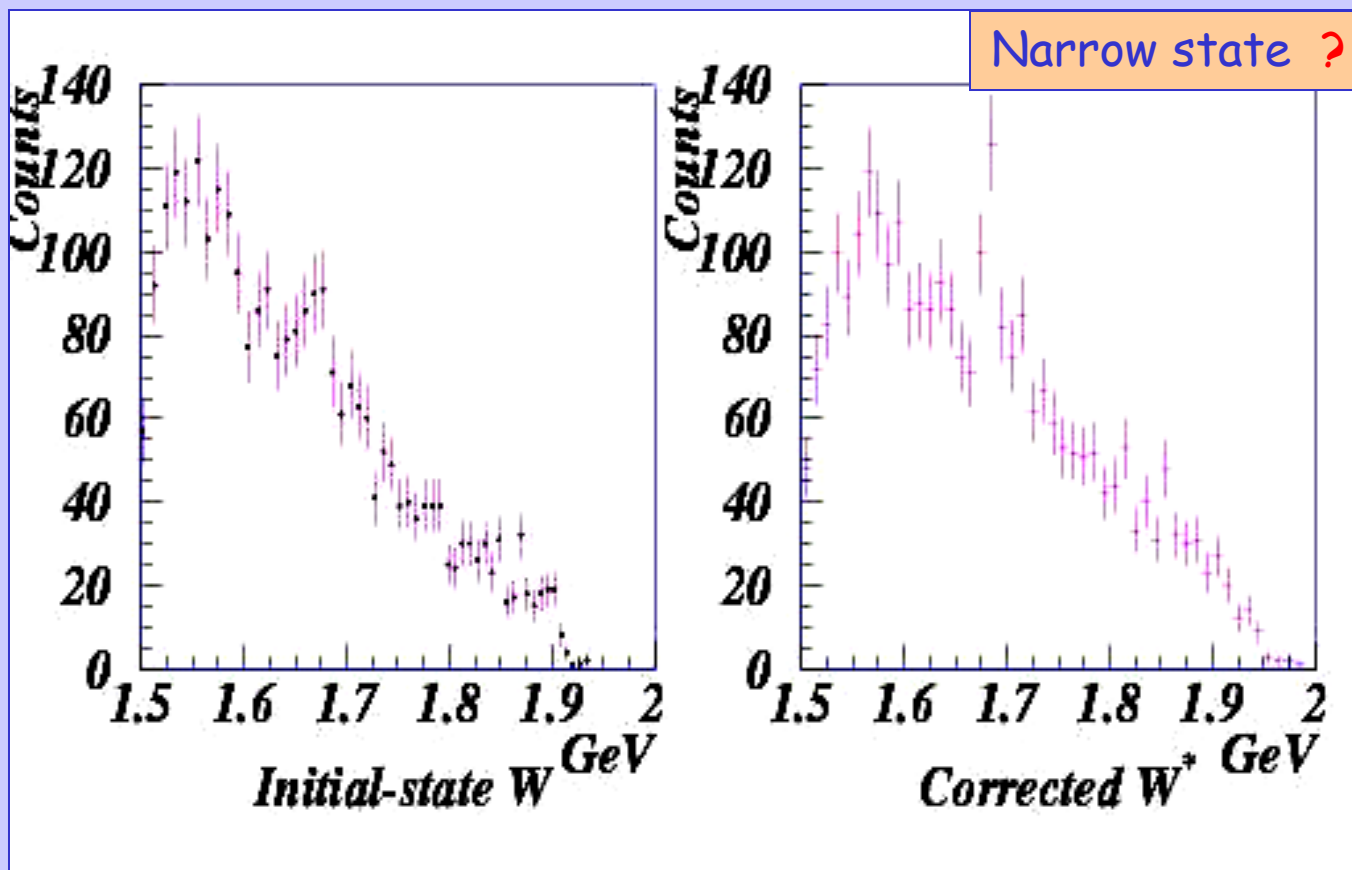
- The SAID soln for the η production off proton scaled by factor 0.6, as has been suggested by previous experiments, fits well the Xsection off the neutron in the region of the $N(1535)D_{15}$ below $W \sim 1.62 \text{ GeV}$
- The sum of the SAID soln, scaled by 0.6, and the simulated contribution of a narrow state ($M = 1.675 \text{ GeV}$, $\Delta W = 10 \text{ MeV}$), fits well Xsection on the neutron up to $W \approx 1.7 \text{ GeV}$!
- This state appears as a wider bump in Xsection due to Fermi motion

GRAAL [V. Kuznetsov, hep-ex/0601002, NSTAR 2005, Oct 2005]

Very preliminary: $\gamma n \rightarrow \eta n$

No correction

Correction for Fermi motion



GRAAL [V. Kuznetsov, Trento, Feb 2004]

Very-very preliminary: $\gamma n \rightarrow K^0 \Lambda$, $K^+ \Sigma^-$

N(1680) ?

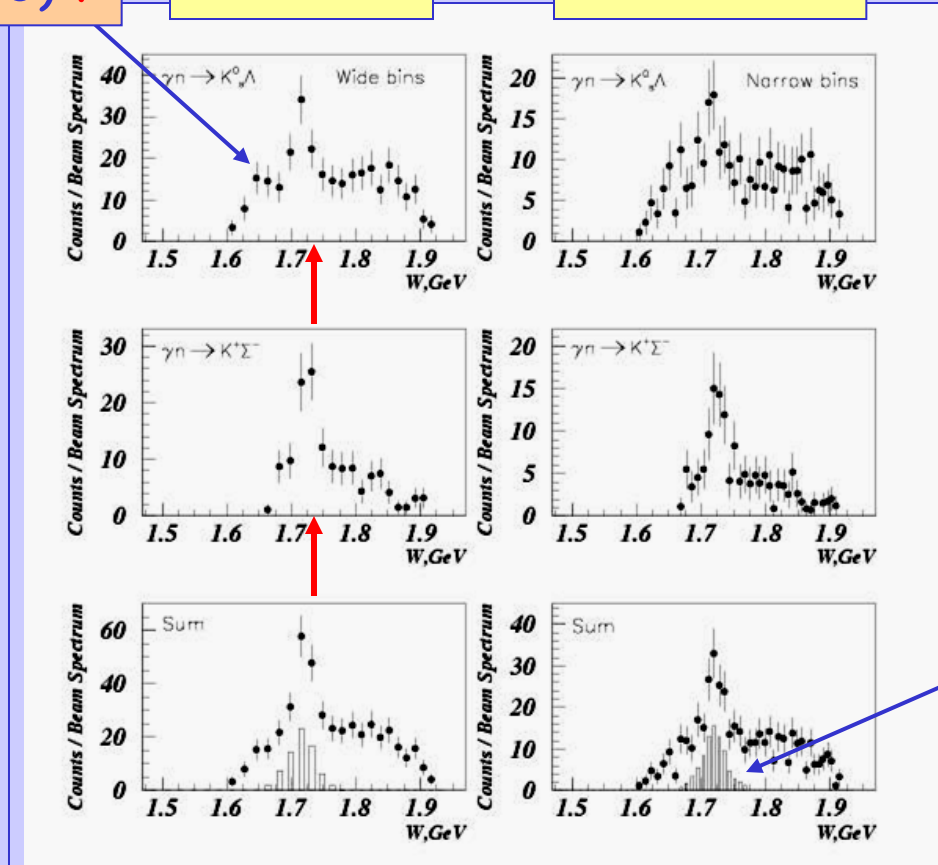
Wide bins

Narrow bins

$\gamma n \rightarrow K^0 \Lambda$

$\gamma n \rightarrow K^+ \Sigma^-$

Sum

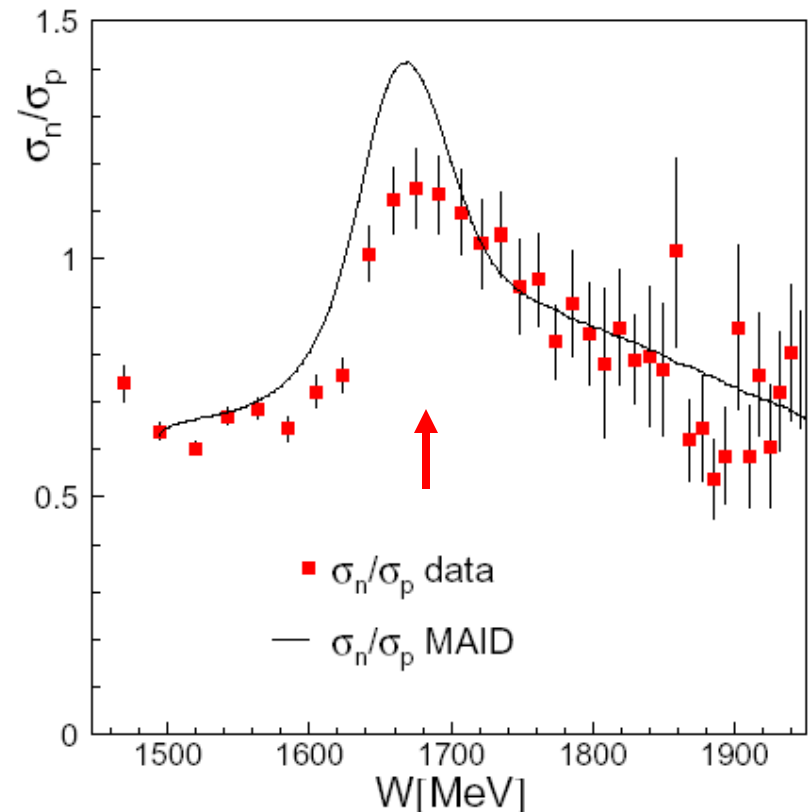
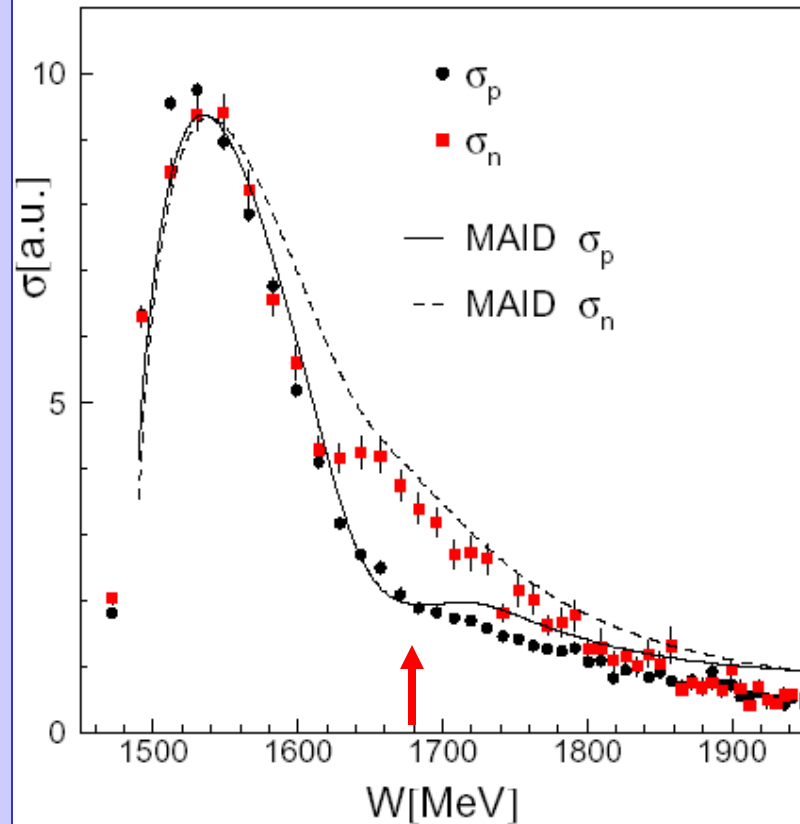


- Quantum numbers of N(1720) are unknown but it could be P_{11} , P_{13} , or less likely S_{11}

N(1720)

CB-ELSA [I. Jaegle, NSTAR 2005, Oct 2005]

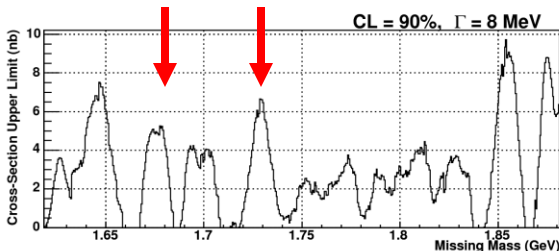
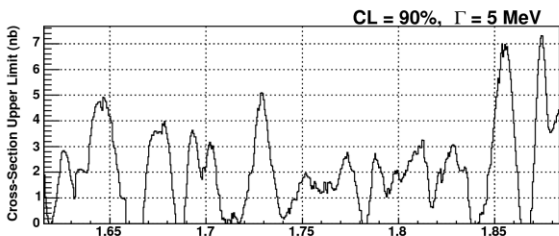
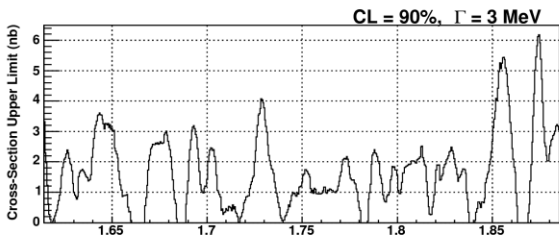
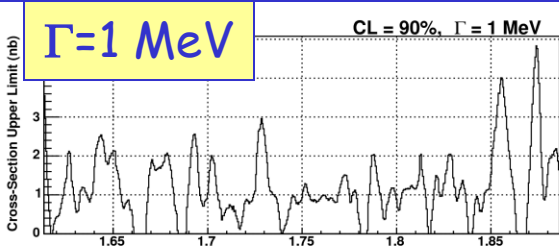
Very preliminary: $\gamma n \rightarrow \eta n$



- Independent CB-ELSA measurements confirm the GRAAL observation

JLab Hall A [B. Wojtsekhowski, Yi Qiang, E-04-012]

Very-very preliminary $H(e, e' \pi^+) X^0$, data taken in May of 2004



$$\Delta\chi^2 = \chi^2(R+B) - \chi^2(B)$$

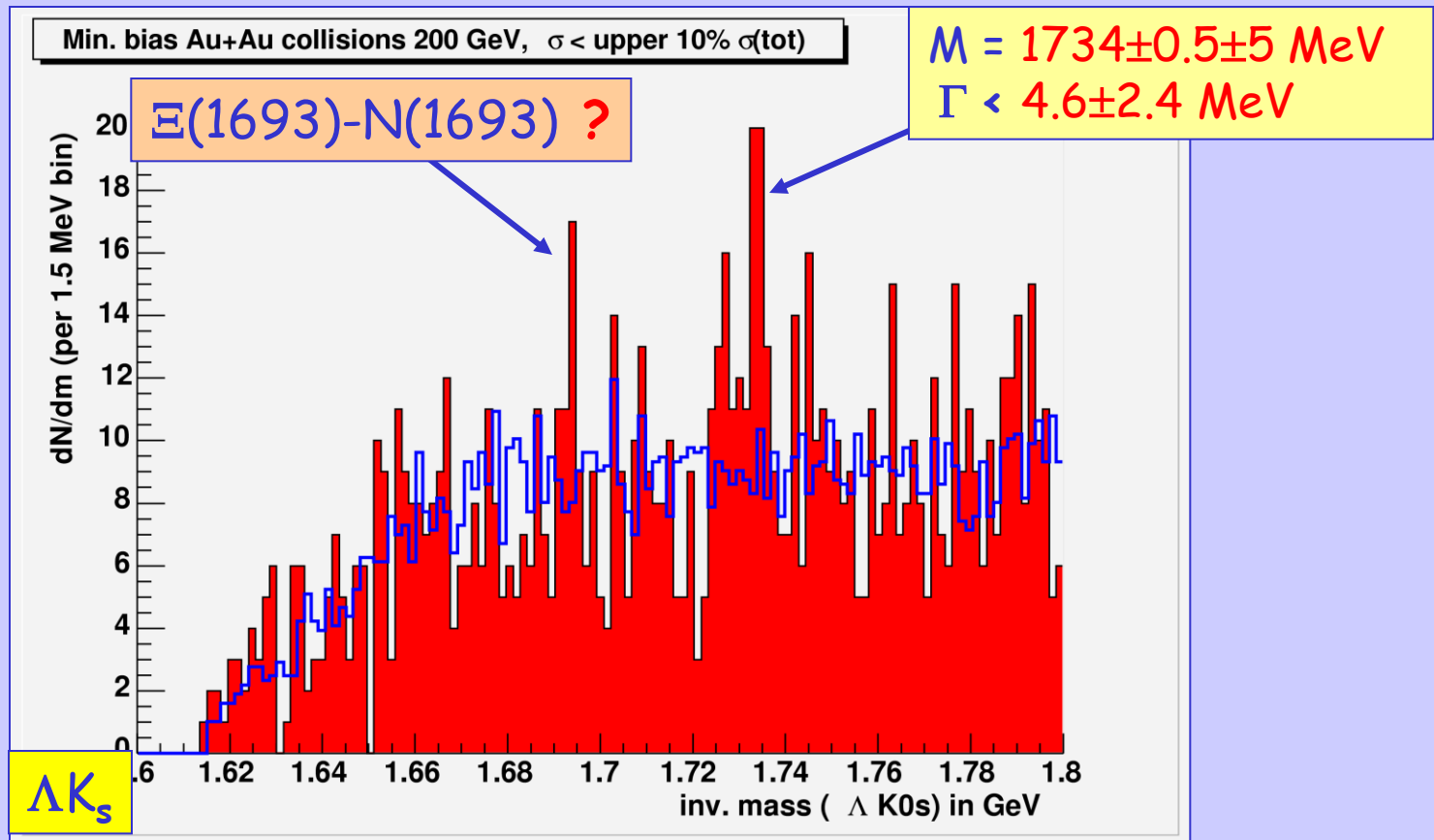
- $E_0 = 5 \text{ GeV}$
- $\theta_{e'} = 6^\circ$
- $\theta_\pi = 0^\circ \quad \Delta\Theta = \pm 2^\circ$
- $\sigma_{MM} = 1.3 \text{ MeV}$

- Signal $N(1680)$ and $N(1730)$ (if any) is small (agrees with expectations)
- The UL result is **strong** for $\Gamma \sim 1 \text{ MeV}$ and becomes **softer** for $\Gamma \sim 10 \text{ MeV}$
- Extraction of $\Gamma_{\pi N}$ needs model assumptions

$\Gamma = 8 \text{ MeV}$

STAR [S. Kabana, hep-ex/0406032, Jamaica, March 2004]

Preliminary: $AuAu \rightarrow \Lambda K_s$



Summary

- Narrowness of Θ^+ required reanalysis of all its flavor partners

We did it for `N(1710)' using modified πN PWA

- If Θ^+ is indeed a narrow state with $\Gamma_{\Theta} \leq 1$ MeV, then other members of the flavor $\overline{10}$ are, most probably, narrow as well

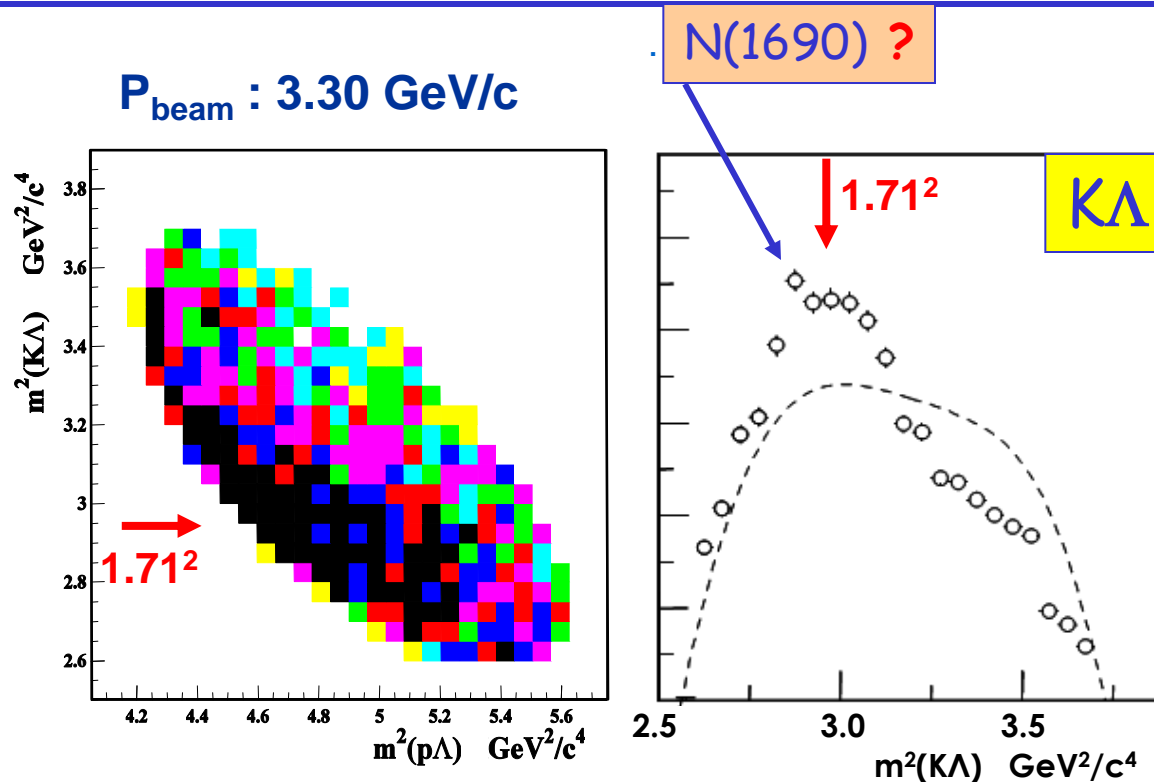
Their properties are sensitive to the structure of mixing which can be rather complicated

- Further measurements/analyses are **necessary** !!

Backup

COSY-TOF [W. Eyrich, Pentaquark 2004, July 2004]

Very preliminary: $pp \rightarrow \Lambda K^+ p$



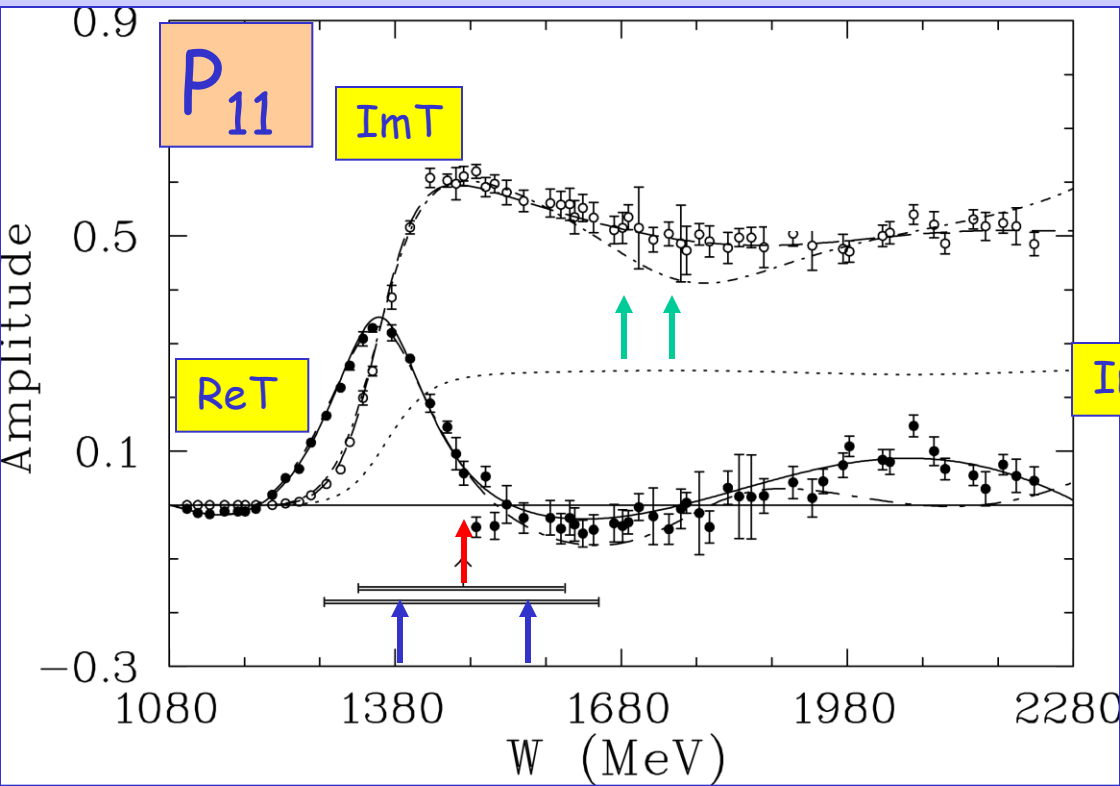
$N^*(1710)$ contributes strongly

Influence of $p\Lambda$ -FSI

In progress: Investigation of Dalitz plots \rightarrow width

P_{11} within πN PWA

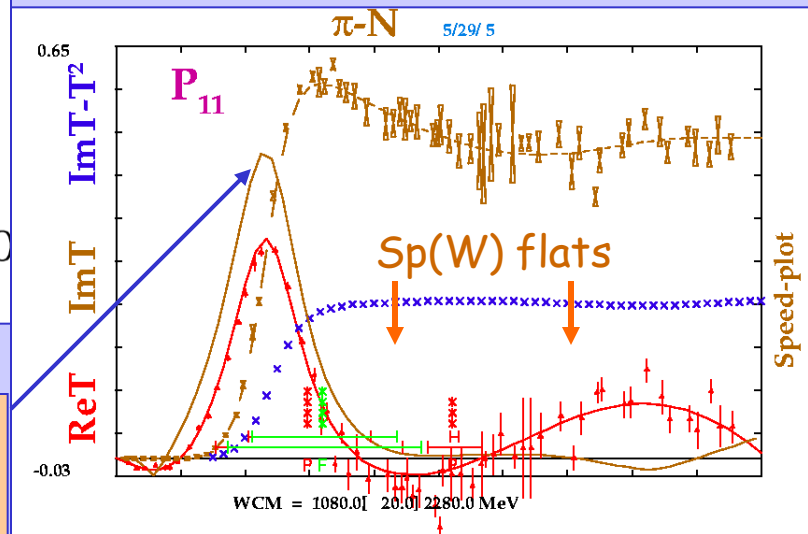
[R. Arndt, W. Briscoe, IS, R. Workman, M. Pavan, Phys Rev C 69, 035213 (2004)]



- $N(1440)P_{11}$ [SAID]
 - BW: $M = 1468.0 \pm 4.5$ MeV
 - $\Gamma/2 = 180 \pm 13$ MeV
 - $X = 0.750 \pm 0.024$
 - Pole: $1357 - i80$ MeV (1st sh)
 - $1385 - i83$ MeV (2nd sh)

$\text{Im}T - T^*T < \text{Im}T$ [unitarity limit]

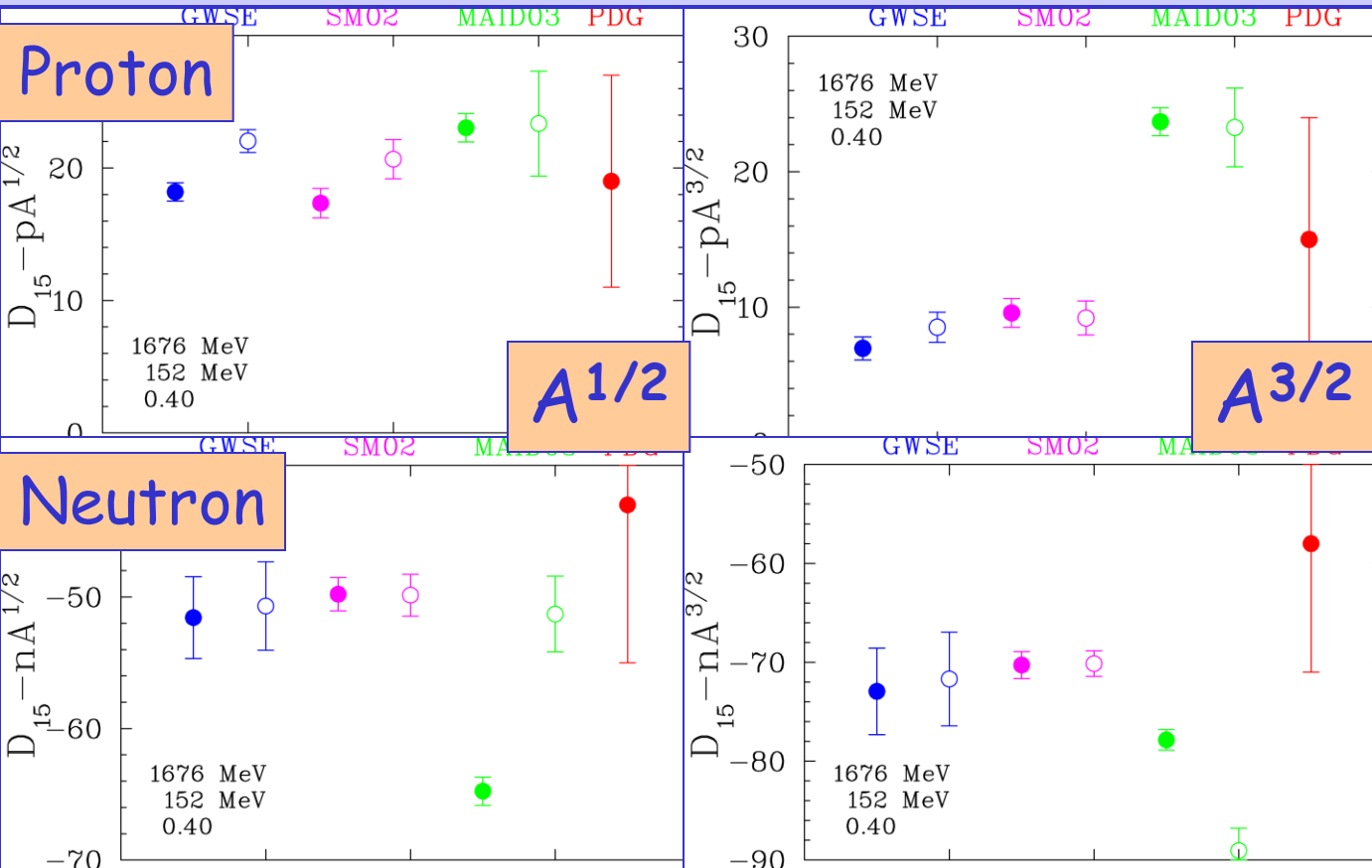
- $\text{Sp}(W) = |dT/dW| \rightarrow$ peak at $W=M(\text{pole})$ at NR $\rightarrow 0$ [G. Hoehler,]



- If Θ^+ does not survive, 'damned' questions revive
-
- 'Why are there no strongly bound exotic states..., like those of two quarks and two antiquarks or four quarks and one antiquark ?'
[H. Lipkin, Phys Lett **45B**, 267 (1973)]
 - '...either these states will be found by experimentalists or our confined, quark-gluon theory of hadrons is as yet lacking in some fundamental, dynamical ingredient which will forbid the existence of these states or elevate them to much higher masses.'
[R. Jaffe and K. Johnson, Phys Lett **60B**, 201 (1976)]

Separation of Res and Nres in $\gamma p \rightarrow \pi N$, $I=1/2, 3/2$

[R. Arndt, W. Briscoe, IS, R. Workman, L. Tiator, in progress]



- A-form: $T = (1 + it_{\pi N})(\text{Born} + A) + R t_{\pi N} + (C + iD)(\text{Im} t_{\pi N} - |t_{\pi N}|^2)$
- C-form: $T = (1 + it_{\pi N})(\text{Born} + A) + R t_{\pi N} e^{i\phi}$

Θ^+ and Φ - What is known

[S. Eidelman *et al* (PDG) Phys Lett B 592, 1 (2004)]

	Experiment	Mass (MeV)	Width (MeV)
$\Theta(1540)^+$	LEPS	1540 ± 10	< 25
	DIANA	1539 ± 2	< 9
	CLAS (d)	1542 ± 5	< 21
	SAPHIR	$1540 \pm 4 \pm 2$	< 25
	ITEP (ν)	1533 ± 5	< 20
	CLAS (p)	1555 ± 10	< 26
	PDG average	1539.2 ± 1.6	-
	GWU	1545	≤ 1
	LBNL	1540	0.9 ± 0.3
$\Phi(1860)$	NA49	1862 ± 2	< 18

Only one pw P_{01} admits the effect at 1540 - 1450 MeV with $\Gamma < 1$ MeV
 [R. Arndt, IS, R. Workman, Phys Rev C 68, 042201 (2003)]

With additional assumption and unknown systematics
 [R. Cahn and G. Trilling, Phys Rev D 69, 011501 (2004)]

The measured mass looks similar to expectation of the χ_{SA}
 [D. Diakonov, V. Petrov, M. Polyakov, Z Phys A 359, 305 (1997)]